PRESSURE DEPENDENCE OF DEFECT EMISSIONS AND THE APPEARANCE OF PRESSURE-INDUCED DEEP CENTERS IN CHALCOPYRITE ALLOYS Ag<sub>x</sub>Cu<sub>1-x</sub>GaS<sub>2</sub>

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**ABSTRACT** 

We present the pressure dependence of the defect emissions in the

chalcopyrite alloy semiconductor Ag<sub>x</sub>Cu<sub>1-x</sub>GaS<sub>2</sub> for values of the alloy

concentration x varying between 0 and 1. A large variation in the pressure

coefficients of the different defect emissions with x was found. In one

alloy concentration x=0.25 deep levels were found to appear under

pressure. Plausible explanations of our results have been proposed.

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Recently alloys of chalcopyrite semiconductors such as

CuGa<sub>x</sub>In<sub>1-x</sub>Se<sub>2</sub> have received attention for improving the performance of solar cells[1]. As a result, the effect of alloying on the electronic band structure and vibrational properties of several chalcopyrite alloy systems have been investigated[2-5]. However, the effect of alloying on the defect properties in chalcopyrite semiconductors is still largely unexplored[3]. It is now well-established that donors in alloys of zincblende-type semiconductors, like AlGaAs, exhibit a shallow-to-deep transformation as a function of both pressure and alloy concentration[6]. In view of the close relation between the chalcopyrite and the zincblende structures, it is interesting to ask whether this phenomenon can also occur in the chalcopyrite semiconductors. In this paper we have studied the pressure dependence of defect photoluminescence (PL) in Ag<sub>x</sub>Cu<sub>1-x</sub>GaS<sub>2</sub> for several alloy compositions. While we did not see shallow-to-deep transformation as a function of alloy concentration, we observed signatures of such transformation as a function of pressure in the alloy Ag<sub>0.25</sub>Cu<sub>0.75</sub>GaS<sub>2</sub>. A lack of reliable band structure calculation in the Ag<sub>x</sub>Cu<sub>1-x</sub>GaS<sub>2</sub> alloys precludes at present a definitive explanation of this result. However, qualitative explanations are proposed.

We have grown alloys of  $Ag_xCu_{1-x}GaS_2$  with x=0.25, 0.5, and 0.75 with the horizontal Bridgman method. The samples were not intentionally doped. They have been characterized by x-ray diffraction and found to be single-phased and crystalline. Details of the properties of these samples

have been presented elsewhere [5] and will not be reproduced here. The high pressure optical experiments were performed with the sample placed inside a standard gasketed diamond-anvil cell using a mixture of methanol and ethanol as the pressure medium. The pressure inside the cell was determined by the ruby fluorescence technique. The entire cell was cooled to 50K in a closed-cycle He refrigerator. The photoluminescence measurement was excited by an Ar<sup>+</sup> laser, analyzed with a double spectrometer and detected with a photon counting system.

In Fig. 1 (a)-(c) we show a series of pressure dependent PL spectra in Ag<sub>x</sub>Cu<sub>1-x</sub>GaS<sub>2</sub> alloys for x=0.25, 0.5 and 0.75. When compared with the corresponding spectra in AgGaS<sub>2</sub> and CuGaS<sub>2</sub> samples grown under similar conditions [5] we notice similarities between the alloy PL spectra and those of AgGaS<sub>2</sub> and CuGaS<sub>2</sub>. In all the samples either one or two defect emission peaks are observed. In Ag<sub>0.75</sub>Cu<sub>0.25</sub>GaS<sub>2</sub> only one emission peak is observed as in CuGaS<sub>2</sub>. On the other hand, the x=0.25 and 0. 5 samples show two PL peaks as in AgGaS<sub>2</sub>. The higher energy peak in both AgGaS<sub>2</sub> and CuGaS<sub>2</sub> has previously been labeled B[5]. It has been attributed to donor-acceptor pair recombination and has been found to follow the band gap as a function of alloy concentration[5]. The lower energy peak, labeled previously as C, has been attributed to deep centers emission and its energy does not follow the band gap as a function of x[5].

In Table 1 we list the pressure coefficients of the PL peaks B and C in  $Ag_xCu_{1-x}GaS_2$  alloys. For comparison purpose we have included the

corresponding pressure coefficients in both AgGaS<sub>2</sub> and CuGaS<sub>2</sub> and also the band gap (E<sub>g</sub>) pressure coefficients measured in separate optical absorption experiments[8]. In the case of Ag<sub>0.25</sub>Cu<sub>0.75</sub>GaS<sub>2</sub> the PL spectra shown in Fig. 1 (a) indicate that both B and C exhibit very striking changes at pressures around 4 GPa. Above 3 GPa. Peak B weakens in intensity and disappears while another peak (which is labeled PB in Fig. 1(a)) appears above 5 GPa. This new peaks is narrower and has completely different pressure dependence from the peak B observed at lower pressures. The peak C seems to split into three narrower peaks around 5 GPa. While they are not very well resolved at low pressure, they become stronger in intensity and become much better resolved at higher pressure. The broken curves in Fig. 1(a) show the three individual peaks while the solid curve is their sum. The experimental points are shown in open circles. These peaks have been labeled PC<sub>1</sub> to PC<sub>3</sub>. One striking characteristics of these new peaks appearing under high pressure is their very small and negative pressure coefficients. The pressure coefficients of these new peaks are also listed in Table 1. Similar tendency of the peak C to show a negative pressure coefficient is found in the other two alloys with x=0.5 and x=0.75. However it is more difficult to decide in these alloys whether there are additional fine structures in peak C.

In the pure compounds we have found that the pressure coefficient of the peak B to be larger than that of the band gap. This remains true in the alloys except for x=0.75. Unlike the case of  $CuGa_xIn_{1-x}Se_2$  whose

band gap pressure coefficient  $dE_g/dP$  varies linearly with x[4], the variation of  $dE_g/dP$  in  $Ag_xCu_{1-x}GaS_2$  with x is highly nonlinear as shown in Table 1. However, the variation of the pressure coefficients of the defect emission peaks with x is even more nonlinear than that of the band gap. The only trend one can deduce from Table 1 is that the pressure coefficients of the defect emission B in the alloys are always *smaller* than those in the pure compounds.

So far there has been very little theoretical studies of the effect of alloying on the band structure and defect properties in chalcopyrite alloys. The only report known to us [3] deals with the alloy system: CuGa<sub>x</sub>In<sub>1</sub>. <sub>x</sub>Se<sub>2</sub>. In comparing the results in this alloy system with our results in Ag<sub>x</sub>Cu<sub>1-x</sub>GaS<sub>2</sub> one has to keep in mind the important differences between the two systems. In the former alloy system, In is replaced by another nontransition metal ion Ga. As pointed out by Wei et al.[3] the main effect of this replacement on the band structure is to change the conduction band. The difference in electronegativity and ionic radii between In and Ga contributes to a rather large bowing parameter of ~0.2 eV in CuGa<sub>x</sub>In<sub>1</sub>. <sub>x</sub>Se<sub>2</sub>[3]. However, the perturbation on the Cu d-bands is minimal. In the Ag<sub>x</sub>Cu<sub>1-x</sub>GaS<sub>2</sub> system, Cu is replaced by another transition metal ion Ag. In addition to their difference in electronegativity and ionic radii there is now significant difference in their d-bands. It is known that the valence bands in the chalcopyrites are strongly affected by the p-d hybridization[9-10]. Thus replacing Cu by Ag is expected to have strong effects on both the valence and conduction bands. One indication of this strong effect is the large bowing parameter of  $\sim 0.8$  eV reported in  $Ag_xCu_{1-x}GaS_2[5]$ . The different dependence of  $dE_g/dP$  on x is obviously another indication.

We have noted previously that the pressure coefficient dE<sub>B</sub>/dP of the defect emission peak B in the chalcopyrite semiconductors AgGaS<sub>2</sub> and CuGaS<sub>2</sub> are larger than the band gap coefficient dE<sub>g</sub>/dP as a result of p-d hybridization in the valence bands[10]. Increasing this hybridization has the effect of decreasing the pressure coefficients since the d-bands tend to be less sensitive to pressure than the p-bands. Perhaps the most dramatic demonstration of this effect is the large drop in both dE<sub>B</sub>/dP and  $dE_g/dP$  when 25% of Ag is added to CuGaS2. An increase in  $dE_g/dP$  is observed when Cu is added to Ag also the effect is less striking. The appearance of the new peaks PB and PC<sub>i</sub> (i=1-3) under pressure in Ag<sub>0.25</sub>Cu<sub>0.75</sub>GaS<sub>2</sub> may also be related to the change in hybridization with alloying. However, the fact that these new peaks have very small and negative pressure coefficients also suggest that the conduction bands are altered in a significant way under pressure so as to change the nature of the donor (or donors) involved in the emission peak. As noted earlier the donor involved in the peak B has been understood to be a shallow donor. However, the donor involved in the peak PB is most likely a deep donor since its pressure dependence is quite different from that of the band gap.

At first sight the behavior of the PL peak B under pressure is rather similar to the conversion of shallow donors in GaAs to deep donors known

as the DX centers under pressure[6]. However this transformation in III-V semiconductors are known to result from the near degeneracy of their conduction band minima under pressure. The I-III-VI $_2$  chalcopyrite semiconductors are more similar to the II-VI zincblende semiconductors (like CdS and ZnSe) whose X conduction band minima are typically too high in energy above the  $\Gamma$  minimum for pressure to induce a cross-over. A more likely scenario is the existence of pressure induced lowering of a resonance deep level into the band gap. Alloying has also been known to facilitate such crossing of resonant deep levels and the  $\Gamma$  conduction band minimum and shallow donor levels. Examples of such crossings include N in  $Ga_xAs_{1-x}P$  [12] and oxygen in GaN[13]. Thus we suggest that deep donor levels resonant with the conduction band may exist in  $Ag_xCu_{1-x}GaS_2$  alloys. These deep donors can be lowered into the bandgap by pressure and this process is facilitated by alloying Cu with Ag.

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 $\label{eq:Table 1} \textbf{Table 1} \mbox{ Pressure coefficient of band gap } (E_g), \mbox{ defect emission peaks B, C,} \\ \mbox{PB, PC}_1, \mbox{PC}_2 \mbox{ and PC}_3 \mbox{ in } Ag_x \mbox{Cu}_{1-x} \mbox{GaS}_2 \mbox{ compared with the pure} \\ \mbox{ compounds } Ag \mbox{GaS}_2 \mbox{ and } \mbox{CuGaS}_2. \mbox{ All units are in } meV/\mbox{GPa}.$ 

Peak	AgGaS <sub>2</sub>	$Ag_{x}Cu_{1-x}GaS_{2}$			CuGaS <sub>2</sub>
		x=0.75	x=0.5	x=0.25	
E <sub>g</sub> (<77	37-40.5	43	36	33	42
K)					
В	58	38	42	38	67
С	41			31	24
PB				~0	
PC <sub>1</sub>				-2	
PC <sub>2</sub>				-4	
PC <sub>3</sub>				-1	

## **Figure Captions**

**Fig. 1** The photoluminescence spectra in the alloy (a) Ag<sub>0.25</sub>Cu<sub>0.75</sub>GaS<sub>2</sub>, (b) Ag<sub>0.5</sub>Cu<sub>0.5</sub>GaS<sub>2</sub> and (c) Ag<sub>0.75</sub>Cu<sub>0.25</sub>GaS<sub>2</sub> at T=50K for several different pressures. In case of the high pressure spectra in (a) the broken curves represents the deconvolution of the peak C into three peaks denoted by the vertical arrows. The solid curve is their sum while the experimental points are shown as open circles.

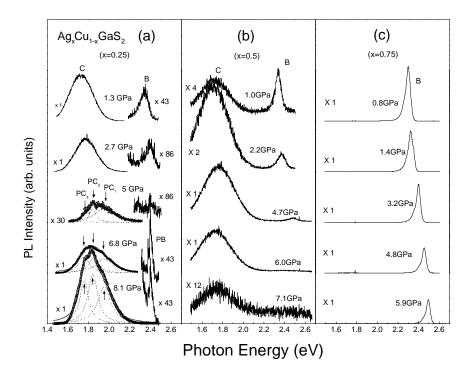


Fig. 1

## **REFERENCES**

- [1] James R. Sites, and Xiaoxiang Liu, Solar Energy Materials & Solar Cells, **41/42**, 373 (1996).
- [2] Zunger, A and Su-Huai Wei, AIP Conference Proceedings, (no.353), (13th NREL Photovoltaics Program Review, Lakewood, CO, USA, 16-19 May 1995.) AIP, 1996. p.155-60.
- [3] S. H. Wei, S. B. Zhang and A. Zunger, Appl. Phys. Lett. **72**, 3199 (1998).
- [4] In-Hwan Choi and Peter Y. Yu, Phys. Stat. Solidi(b)**211**, 51(1999).
- [5] In-Hwan Choi, Sung-Hwan Eom and P. Y. Yu, J. Appl. Phys. **87**, 3815 (2000).
- [6] See, for example, review paper by Ming-fu Li and Peter Y. Yu, in Semiconductors and Semimetals, Vol. 54a, edited by T. Suski and W. Paul(Academic Press, New York, 1998) p.457.
- [7] Preliminary results have been presented in In-Hwan Choi, Sung-Hwan Eom and P. Y. Yu, Phys. Rev.B **61**, 4689 (2000).
- [8] In-Hwan Choi and Peter Y. Yu, Phys. Stat. Solidi(b)211, 143 (1999).
- [9] J. E. Jaffe and A. Zunger, Phys. Rev. B **29**, 1882 (1884).
- [10] In-Hwan Choi and Peter Y. Yu, Phys Rev B **55**, 9642 (1997).
- [12] D. J. Wolford, J. A. Bradley, K. Fry, J. Thompson, in *Physics of Semiconductors 1984*, ed. By J. D. Chadi and W. A. Harrison (Springer, New York, 1984) p.627.

[13] C. Wetzel, T. Suski, J. W. Ager, E. R. Weber, E. E. Haller, S. Fischer, B. K. Meyer, R.J. Molnar, and P. Perlin, Phy. Rev. Lett. 78, 3923 (1997).